# Rotational Mixing and Lithium Depletion

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### **Pre-MS Lithium Depletion**

- Mass and composition trend predicted (Iben 1965)
- Long-standing problem: lithium dispersion in cool stars
  - Dispersion is real, not induced by activity (King et al. 2009)
  - Not from mixing

Lithium in the Pleiades (Soderblom et al. 1993) with varying spot covering factors



### **Culprit: Starspots**

**Pleiades Lithium Revisited** 



Implication: range of radii in pre-MS stars will cause a range of depletion rates

 Anomaly tied to pre-MS activity and rotation



Teff (K)

#### Where does rotation come from?

Initial Angular Momentum Mass assembly rate Interactions with accretion disk Impacted by - Environment Planet formation



Rebull et al. 2006: Correlation of IR excess with rotation period in ONC

## **Angular Momentum Evolution**

#### Magnetized solar-like winds

- more efficient for rapid rotation
- Saturate at massdependent threshold
- Angular Momentum Transport and Mixing
  - Core eventually coupled to surface; three classes of mechanisms competitive

Tinker et al. 2002: ONC->Pleiades->Hyades



Fig. 1.— The cumulative distribution of  $v \sin i$  for the projected sample of Pleiades stars is shown for 4 different saturation thresholds. The projected Pleiades is the solid line and the observed Hyades distribution is the dotted line.

## **Rotational Mixing and Lithium**

#### Two mechanisms for inducing:

- Shears
- Meridional circulation
- Inhibited by mu gradients
  - Can interact with diffusion
- Key predictions:
  - Rate of depletion declines as stars spin down
  - Different rotation histories
  - -> Different depletion histories

# Strong Arguments for Rotational Mixing: Timing and Dispersion



**Fig. 8.** log n(Li) as a function of mass for Be 32 (open circles), the Hyades (asterisks), M 67 (filled triangles), and NGC 188 (filled circles). The horizontal lines delimits the range covered by Pop. II stars considering the lowest and highest values of the plateau.

Sestito & Randich 2005

Randich et al.

2009: Be 32



# The Problem

Standard and diffusive models are rigorous but incomplete Nonstandard models are more complete, but not rigorous

# Convergence in Rotation Rates of Solar Analogs is Clearly Seen



Interesting Coupling Timescale Rapid rotators are strongly coupled Slow rotator spindown is strongly inconsistent with SB rotation  $-\tau \sim 50 \text{ Myr}$ 

> Denissenkov et al. 2009 (astro-ph/0911.1121)



### Implications for Lithium Depletion

- Internal rotation convergences on a timescale comparable to that for surface rotation
  - Dispersion is generated early in the lifetime of a star
- Predicted magnitude of dispersion is therefore sensitive to
  - Angular momentum loss
  - Accretion disk properties

### Interaction of Rotation and Settling

- Required to understand helioseismology (Richard et al. 1988)
- Stalling of lithium depletion in old open clusters (Randich & Sestito 2005)
- Much recent evidence for settling in globular cluster stars (Korn et al. 2007)
- Diffusion expected late, rotational mixing expected early



# **On Halo Lithium Depletion**

V(Li) = 12 + log(Li/H)

0

6500

- Basic issue: Pop II pattern does not look like the Pop I pattern
- Taken as evidence that the Pop II stars are undepleted
  - Very difficult to understand from stellar interiors point of view
     Halo field
     Chorburn 1994
- Diffusion (late ages) can add to mixing (early)
- Major uncertainty: rotational properties of halo stars (dispersion/depletion) relationship
- Testable in halo tidally synchronized binaries

M67 (Jones et al. 1999)



6000

5500

T. (K)

## Binary Interactions & Ultra-Depleted Stars

- Persistent population of highly Li depleted halo stars (e.g. Thorburn 1994)
- Blue stragglers are highly depleted (Ryan et al. 2002)
- Fraction is consistent with predictions from blue straggler counts (Andronov et al. 2006)



Fraction of sub-turnoff merger products (Andronov, Pinsonneault & Terndrup 2006)

There is something fascinating about science. One gets such wholesale returns of conjecture out of such a trifling investment of fact

Cluster vs. field: differences in disk lifetimes?
Planet host vs. no planet host: planet formation impacts disk

evolution?

